

Bayesian uncertainty analysis: linking simulators and climate

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Climate simulators

Much of climate science falls into the general area of the analysis of complex physical systems by use of computer simulators. Here is a framing quote for the area, from Jon Lockley, manager of the Oxford Supercomputing Centre [from the BBC web-site]

“Whenever possible, everything is done in a supercomputer.

Look at Formula One - it's getting rid of all of its wind tunnels and replacing them with supercomputers. It's the same in the aerospace industry as well.

It means you can all the modelling in the supercomputer and then do just one real world test.”

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Great resource: the Managing Uncertainty in Complex Models web-site

<http://www.mucm.ac.uk/> (for references, papers, toolkit, etc.)

[MUCM is a consortium of U. of Aston, Durham, LSE, Sheffield, Southampton - with Basic Technology funding. Now mutating into MUCM community.]

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This talk is concerned with a general overview of some key issues in this area.

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Galaxy formation

The study of the development of the Universe is carried out by using a Galaxy formation simulator.

The aim is scientific - to gain information about the physical processes underlying the Universe.

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- (viii) **multi-model uncertainty** (usually we have not one but many models related to the physical system)
- (ix) **decision uncertainty** (to use the model to influence real world outcomes, we need to relate things in the world that we can influence to inputs to the simulator and through outputs to actual impacts. These links are uncertain.)

RAPID-WATCH (a NERC funding call)

What are the implications of RAPID-WATCH observing system data and other recent observations for estimates of the risk due to rapid change in the MOC? In this context risk is taken to mean the probability of rapid change in the MOC and the consequent impact on climate (affecting temperatures, precipitation, sea level, for example). This project must:

- * contribute to the MOC observing system assessment in 2011;
- * investigate how observations of the MOC can be used to constrain estimates of the probability of rapid MOC change, including magnitude and rate of change;
- * make sound statistical inferences about the real climate system from model simulations and observations;
- * investigate the dependence of model uncertainty on such factors as changes of resolution;
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However, the truly disappointing thing is that, almost always, stated scientific probabilities are not even the judgements of any individual.

Nor are these probabilities given any other meaning (apart from as quantities in idealised models).

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We shall describe the Bayesian approach to these problems, in which all uncertainties are treated probabilistically and combined by rules of Bayesian analysis.

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The statistical approach based around expectation is termed Bayes linear analysis, based around these updating equations for mean and variance:

$$\begin{aligned} E_z[y] &= E(y) + \text{Cov}(y, z)\text{Var}(z)^{-1}(z - E(z)), \\ \text{Var}_z[y] &= \text{Var}(y) - \text{Cov}(y, z)\text{Var}(z)^{-1}\text{Cov}(z, y) \end{aligned}$$

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The examples that we will describe use Bayes linear methods.

(There are natural (but much more complicated) probabilistic counterparts.)

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We use the emulator either to provide a full joint probabilistic description of all of the function values (full Bayes) or to assess expectations variances and covariances for pairs of function values (Bayes linear).

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We need careful (multi-output) experimental design to choose informative model evaluations, and detailed diagnostics to check emulator validity.

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lets us adjust the prior emulator to an appropriate posterior emulator for $f_i(x)$. [This approach exploits the heuristic that we need many more function evaluations to identify the qualitative form of the model (i.e. choose appropriate forms $g_{ij}(x)$, etc) than to assess the quantitative form of all of the terms in the model - particularly if we fit meaningful regression components.]

Illustration from RAPID (thanks to Danny Williamson)

One of the main aims of the RAPIT programme is to assess the risk of shutdown of the AMOC (Atlantic Meridional Overturning Circulation) which transports heat from the tropics to Northern Europe and how this risk depends on the future emissions scenario for CO₂.

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We had access to some runs of FAMOUS (a lower resolution model), which consisted of 6 scenarios for future CO₂ forcing, and between 40 and 80 runs of FAMOUS under each scenario, with different parameter choices.

[And very little time to do the analysis.]

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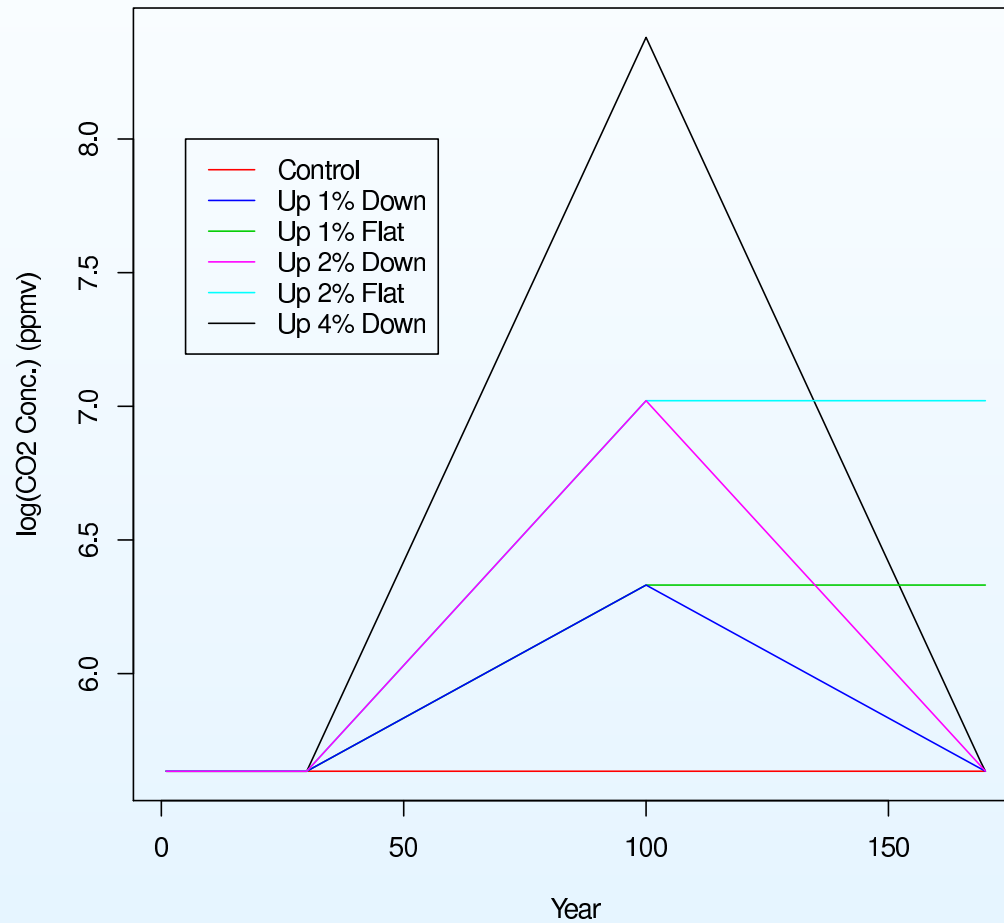
Interest concerned aspects such as:

the value and location of the smoothed minimum of the series

the amount that AMOC responds to CO₂ forcing and recovers if CO₂ forcing is reduced.

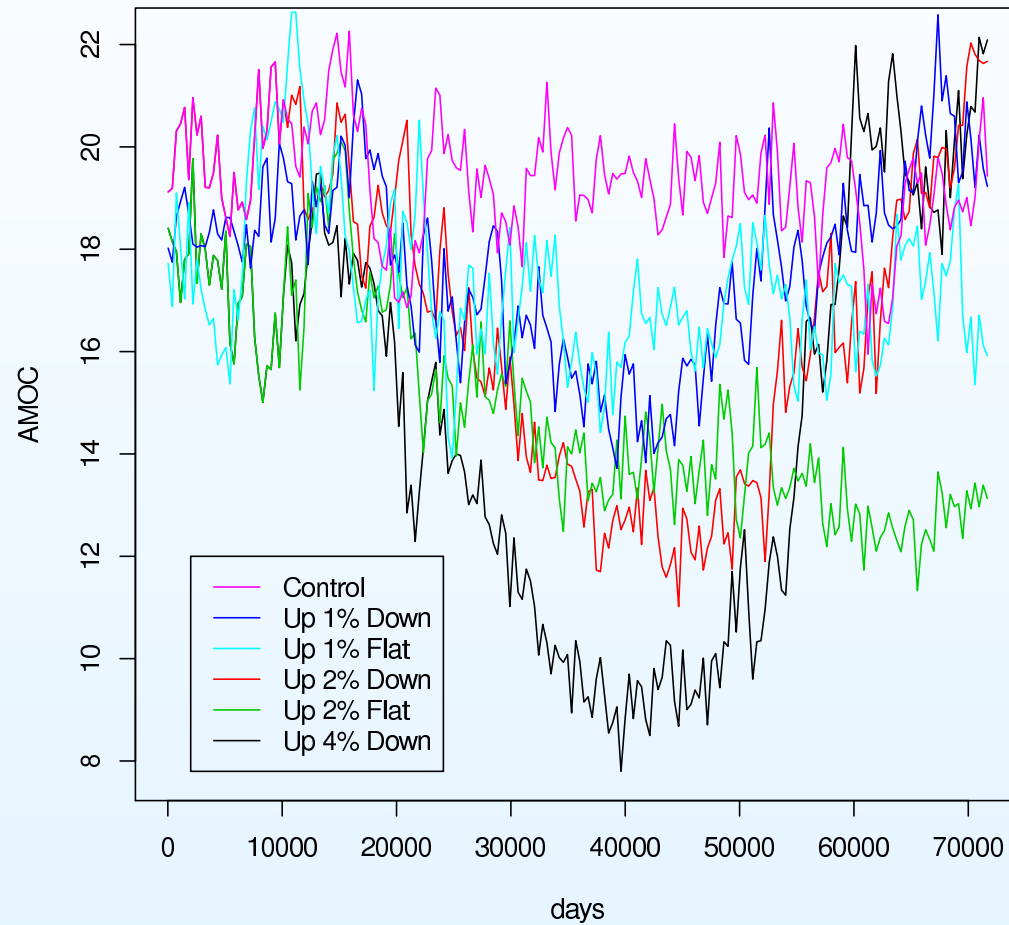
CO2 Scenarios

Log CO2 concentration trajectories



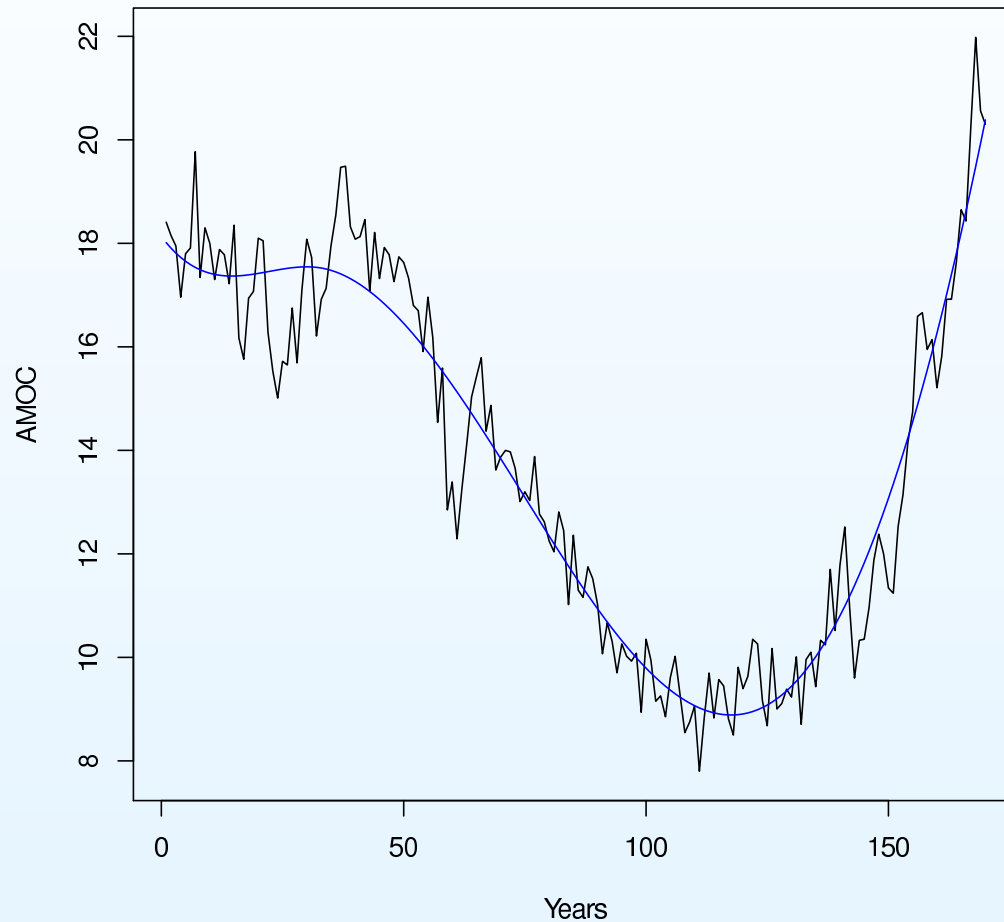
Famous Scenarios

FAMOUS AMOC Scenarios



Smoothing

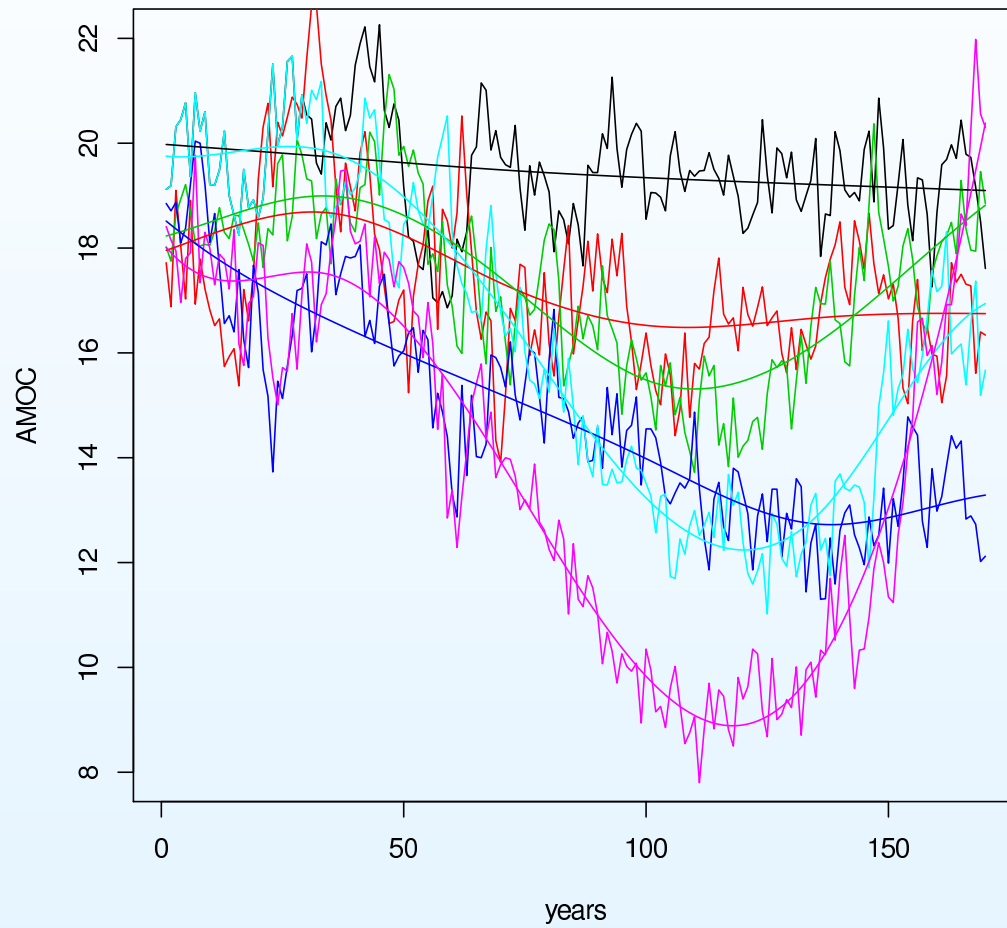
AMOC Up 4% down



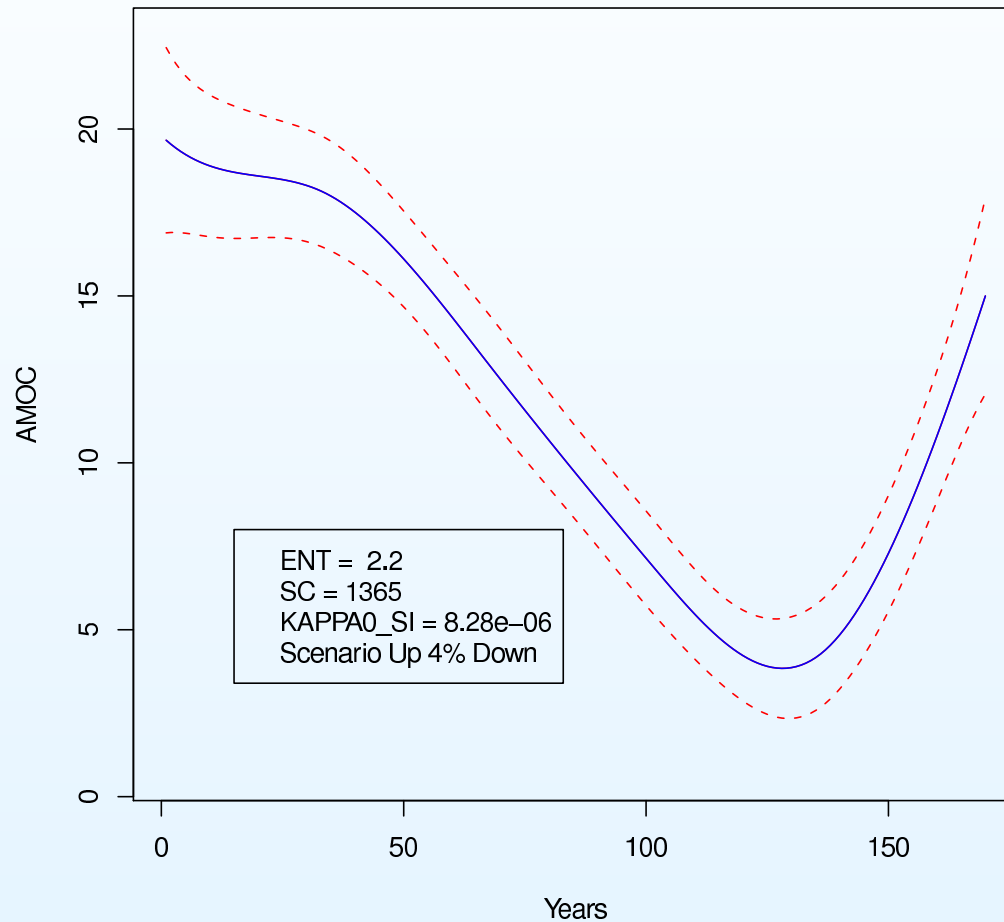
Smooth by fitting splines $f^s(x, t) = \sum_j c_j(x) B_j(t)$ where $B_j(t)$ are basis functions over t and $c_j(x)$ are chosen to give the 'best' smooth fit to the time series.

Smoothing

Splines for each scenario



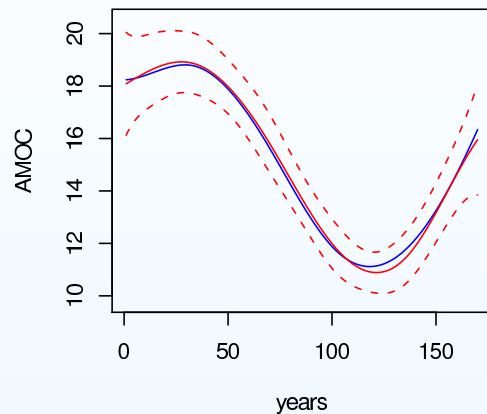
FAMOUS Emulator



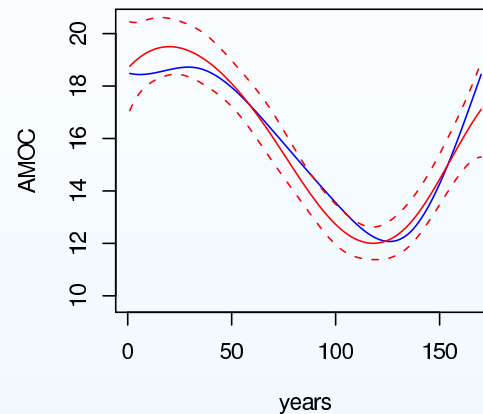
Emulate f^s by emulating each coefficient $c_j(x)$ in $f^s(x, t) = \sum_j c_j(x) B_j(t)$
(separately for each CO2 scenario)

Diagnostics (leave one out)

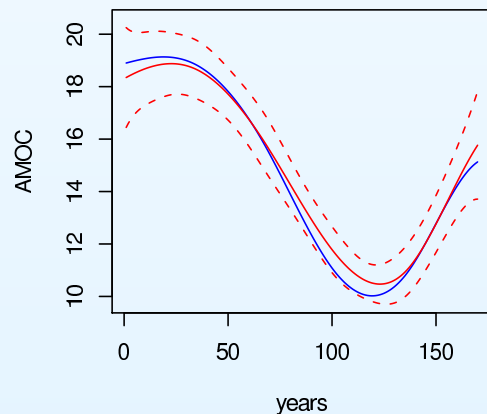
LOO plot for data point 2



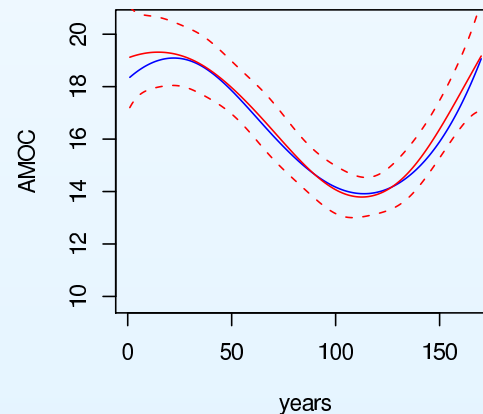
LOO plot for data point 6



LOO plot for data point 7



LOO plot for data point 9



Test approach by building emulators leaving out each observed run in turn, and checking whether the run falls within the stated uncertainty limits.

Emulating HadCM3

We now have an emulator for the smoothed version of FAMOUS, for each of the 6 CO₂ scenarios. Next steps

[1] Extend the FAMOUS emulator across all choices of CO₂ scenario.

[We do this using fast geometric arguments, exploiting the speed of working in inner product spaces. For example, we have a different covariance matrix for local variation at each of 6 CO₂ scenarios. We extend this specification to all possible CO₂ scenarios by identifying each covariance matrix as an element of an appropriate inner product space, and adjusting beliefs over covariance matrix space by projection.]

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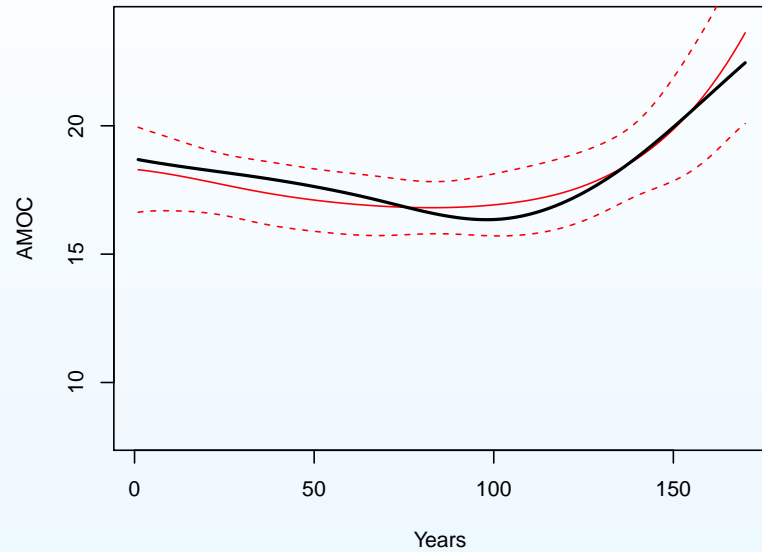
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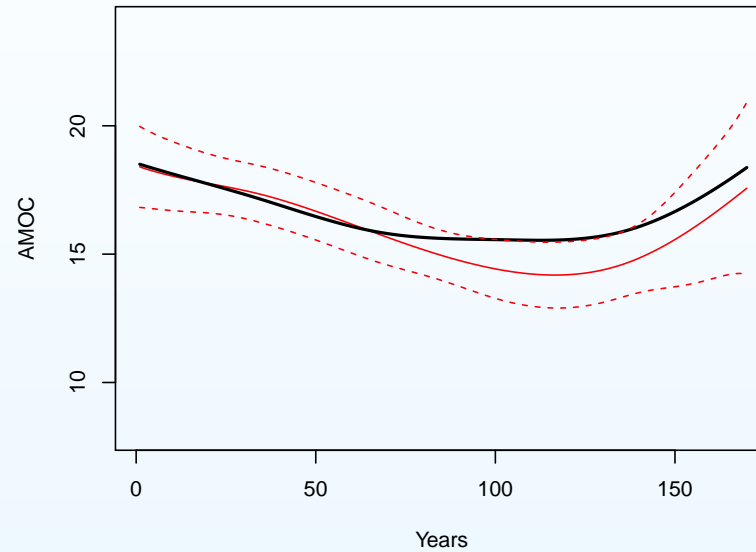
[4] Diagnostic checking, tuning etc.

Emulating HadCM3:diagnostics

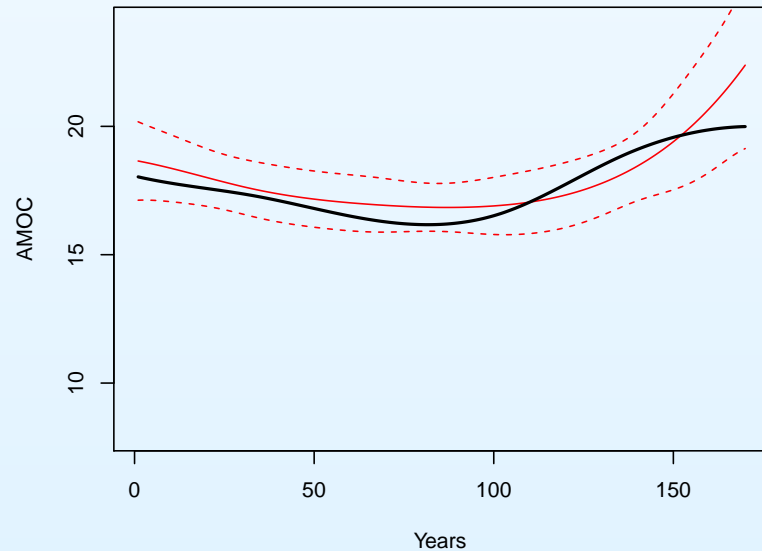
LOO plot for data point 3



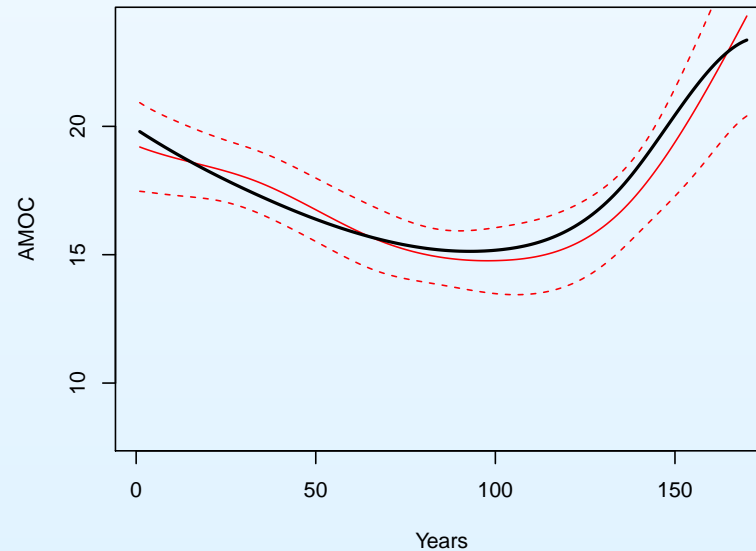
LOO plot for data point 10



LOO plot for data point 13

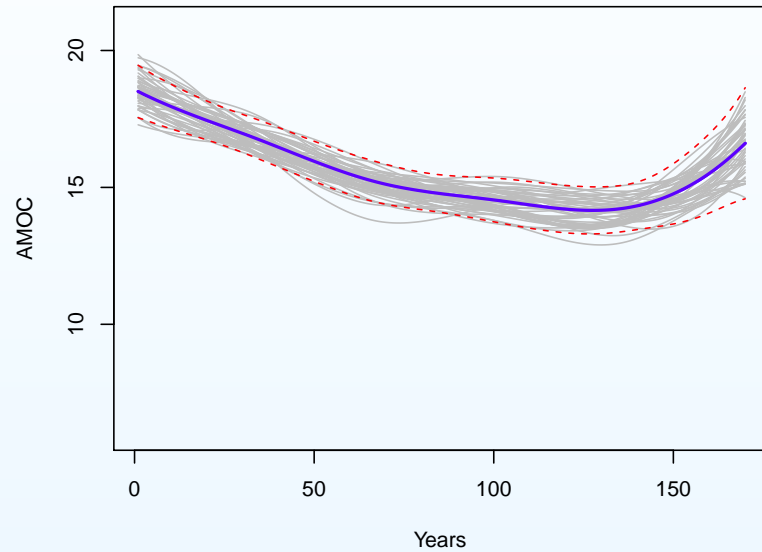


LOO plot for data point 15

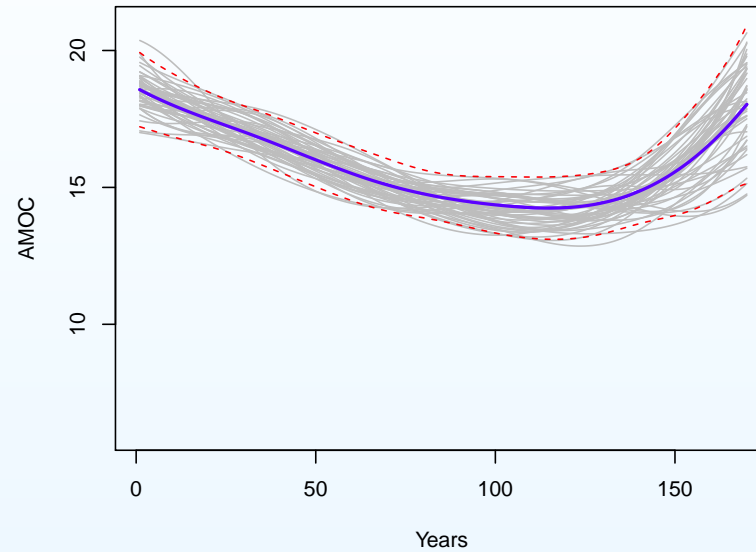


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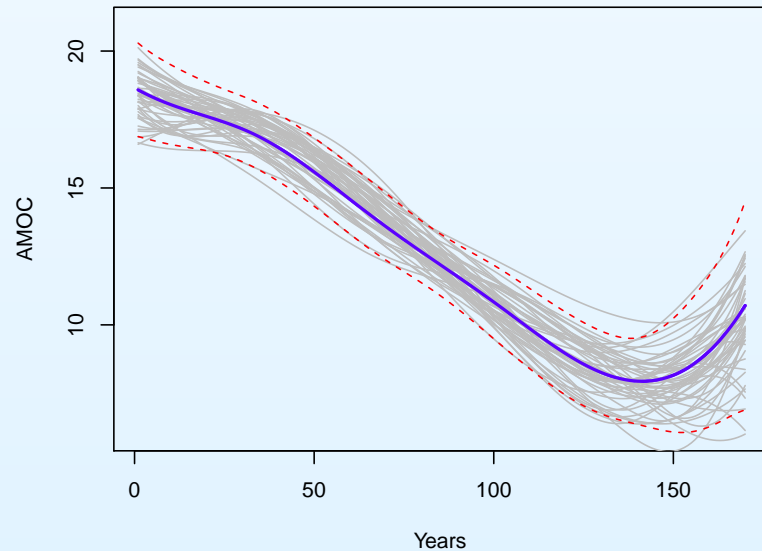
HadCM3 Emulator Up 1.5% Down 0.5%



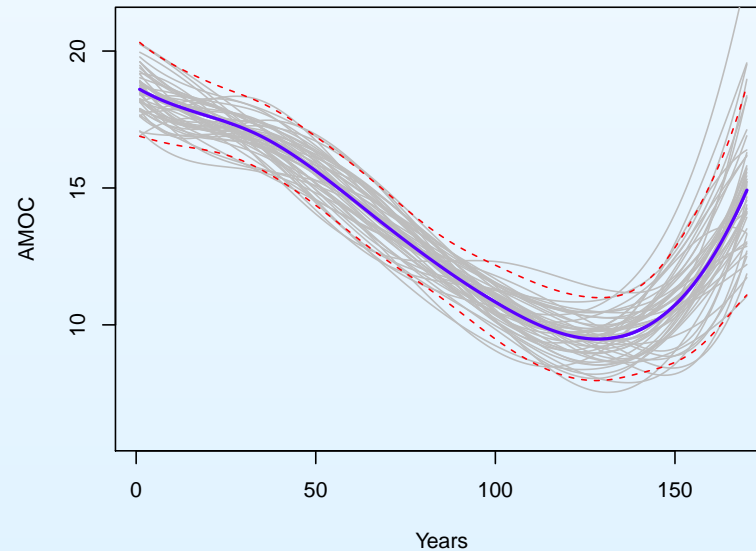
HadCM3 Emulator Up 1.5% Down 0.9%



HadCM3 Emulator Up 2.5% Down 0.5%



HadCM3 Emulator Up 2.5% Down 1.5%



Simulators for Oil Reservoirs

An oil reservoir is an underground region of porous rock which contains oil and/or gas. The hydrocarbons are trapped above by a layer of impermeable rock and below by a body of water, thus creating the reservoir. The oil and gas are pumped out of the reservoir and fluids are pumped into the reservoir (to boost production).

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The simulator outputs comprise the behaviour of the various wells and injectors in the reservoir. Output is time series on features such as bottom hole pressures, oil, gas and water rates.

A reservoir example: (thanks to Jonathan Cumming)

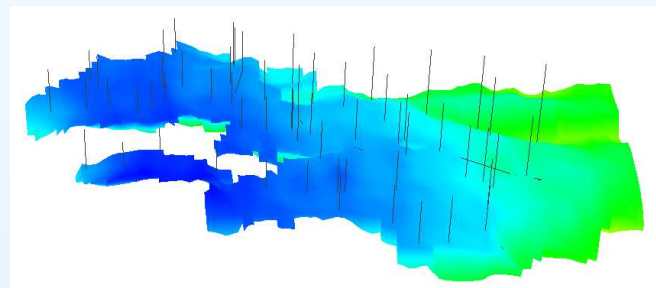
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Inputs Field multipliers for porosity (ϕ), permeabilities (k_x, k_z), critical saturation (crw), and aquifer properties (A_p, A_h)

Outputs Oil production rate for a 3-year period, for the 10 production wells active in that period. 4-month averages over the time series



Coarse and Accurate Emulators

The computer model is expensive to evaluate, so we use 'coarse' model, F^c (by coarsening vertical gridding by factor of 10), to capture qualitative features of F . F^c is substantially faster, allowing 1000 model runs (in a Latin Hypercube over the input parameters).

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For chosen outputs from F^c , we construct emulators f_i^c , from these runs of form

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We now make 20 runs of F , based on a careful choice of small design. We (Bayes linear) update our emulator f for F , supposing the full model emulator to have the form

$$f_i(\mathbf{x}) = \mathbf{g}_i(\mathbf{x}_{[i]})^T \beta_i + w_i^c(\mathbf{x})\beta_{w_i} + w_i^a(\mathbf{x})$$

(linked via equations relating corresponding pairs of coefficients, etc.)

Design considerations

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Typically, the active input sets will be different for different outputs. However, these active sets will not be disjoint.

Border block designs

To deal with the interlocking sets of active variables, we must construct **border-block** designs. eg. if

$$x_{[1]} = \{x_1, x_2, x_3, x_4\}, x_{[2]} = \{x_1, x_2, x_5, x_6\},$$

Then the border is $\{x_1, x_2\}$ and the blocks are $\{x_3, x_4\}$, and $\{x_5, x_6\}$

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How do we identify the border/block separation?

Many runs of fast simulator

These also suggest the types of emulators that we should design for, in terms of natural scales for inputs, key interactions and so forth.

Emulation Summaries

Well	Time	$x_{[i]}$	No. Model Terms	Coarse Simulator R^2	Accurate Simulator \tilde{R}^2
B2	4	ϕ, crw, A_p	9	0.886	0.951
B2	8	ϕ, crw, A_p	7	0.959	0.958
B2	12	ϕ, crw, A_p	10	0.978	0.995
B2	16	ϕ, crw, k_z	7	0.970	0.995
B2	20	ϕ, crw, k_x	11	0.967	0.986
B2	24	ϕ, crw, k_x	10	0.970	0.970
B2	28	ϕ, crw, k_x	10	0.975	0.981
B2	32	ϕ, crw, k_x	11	0.980	0.951
B2	36	ϕ, crw, k_x	11	0.983	0.967

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If the data is informative for the parameter space, then $C(z)$ will typically form a tiny percentage of the original parameter space, so that even if we do wish to calibrate the model, history matching is a useful prior step.

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The implausibility calculation can be performed univariately, or by multivariate calculation over sub-vectors. The implausibilities are then combined, such as by using $I_M(x) = \max_i I_{(i)}(x)$, and can then be used to identify regions of x with large $I_M(x)$ as implausible, i.e. unlikely to be good choices for x^* .

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This process is a form of iterative global search aimed at finding all choices of x^* which would give good fits to historical data.

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We introduce an additional point 1 year beyond the end of the previous series as a value to be forecast (actually, we do have real world observations of this value, to reality check our forecasts).

Forecasting

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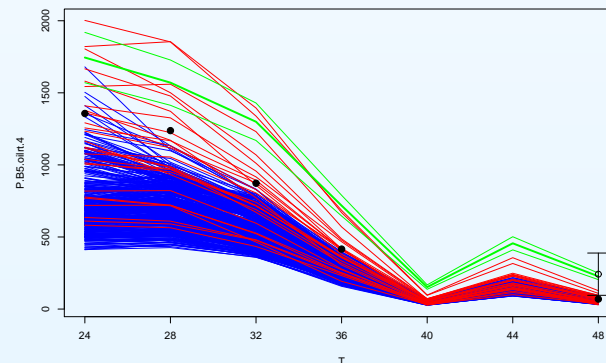
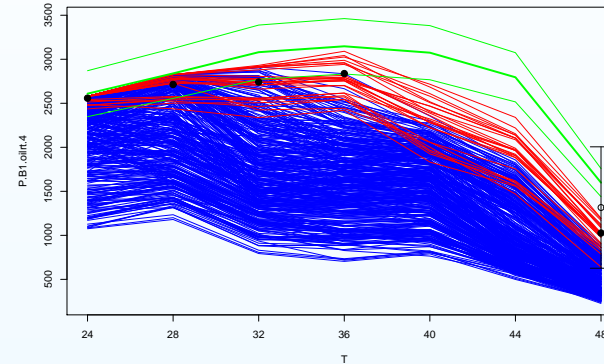
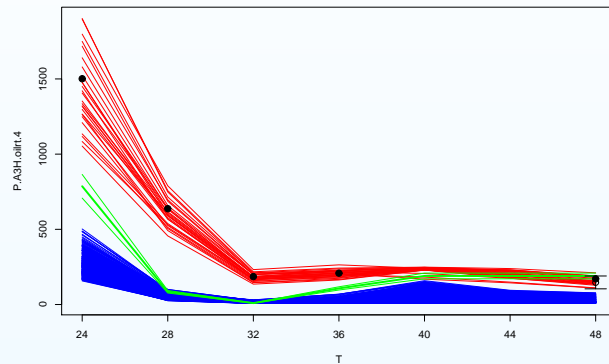
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This analysis is tractable even for large systems with complex descriptions of structural discrepancy, so allows us to vary decision inputs to achieve “optimal control” for the system.

Forecasting Results



Green lines indicate data z with error bounds of $2sd(e)$.

Red and blue lines represent the range of the runs of $F(x)$ and $F^c(x)$

Solid black dots correspond to $E(F^*)$.

The system forecast is indicated by a hollow circle with attached error bars.

(Note how system forecasts correct simulator forecast towards data.)

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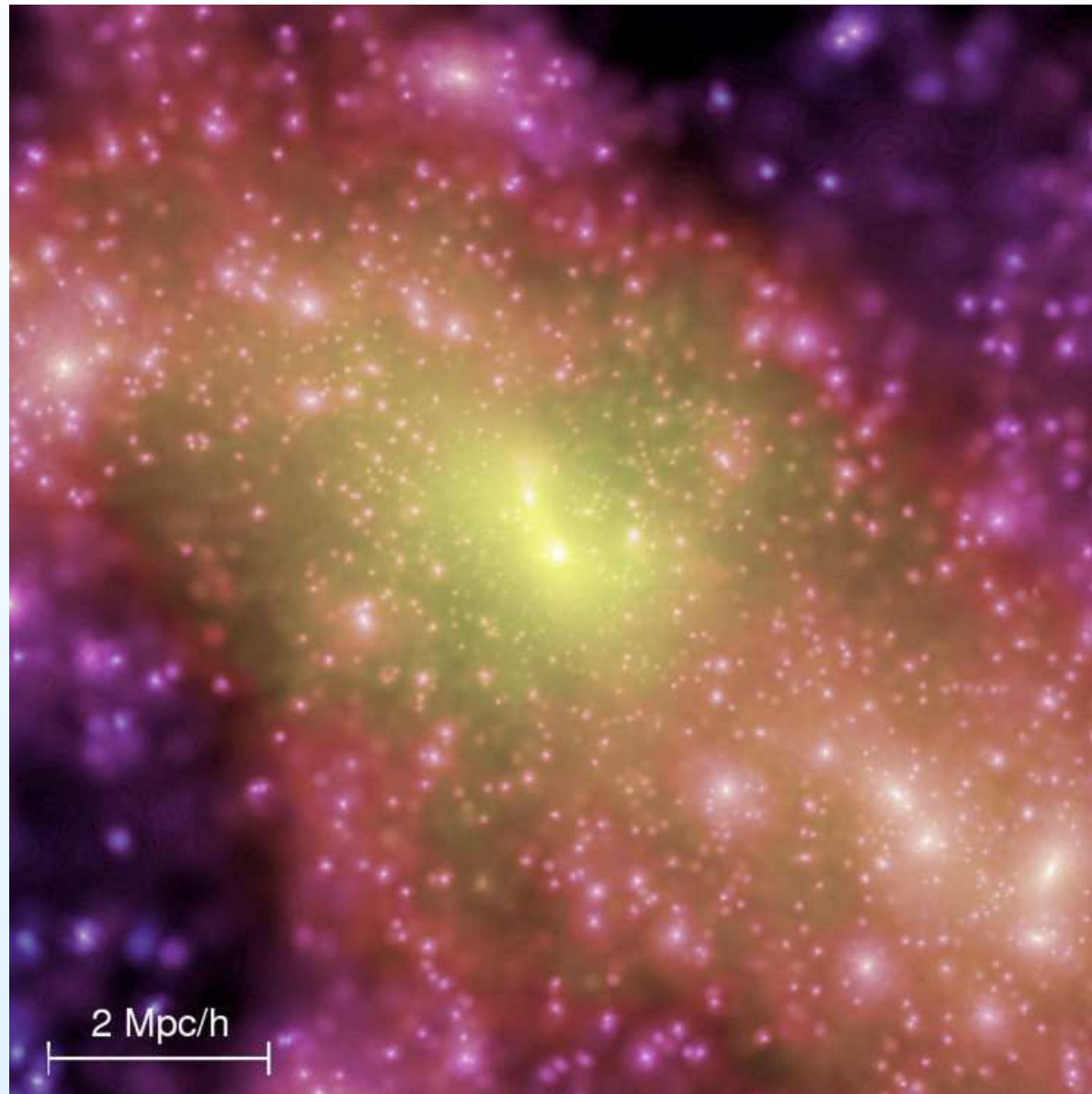
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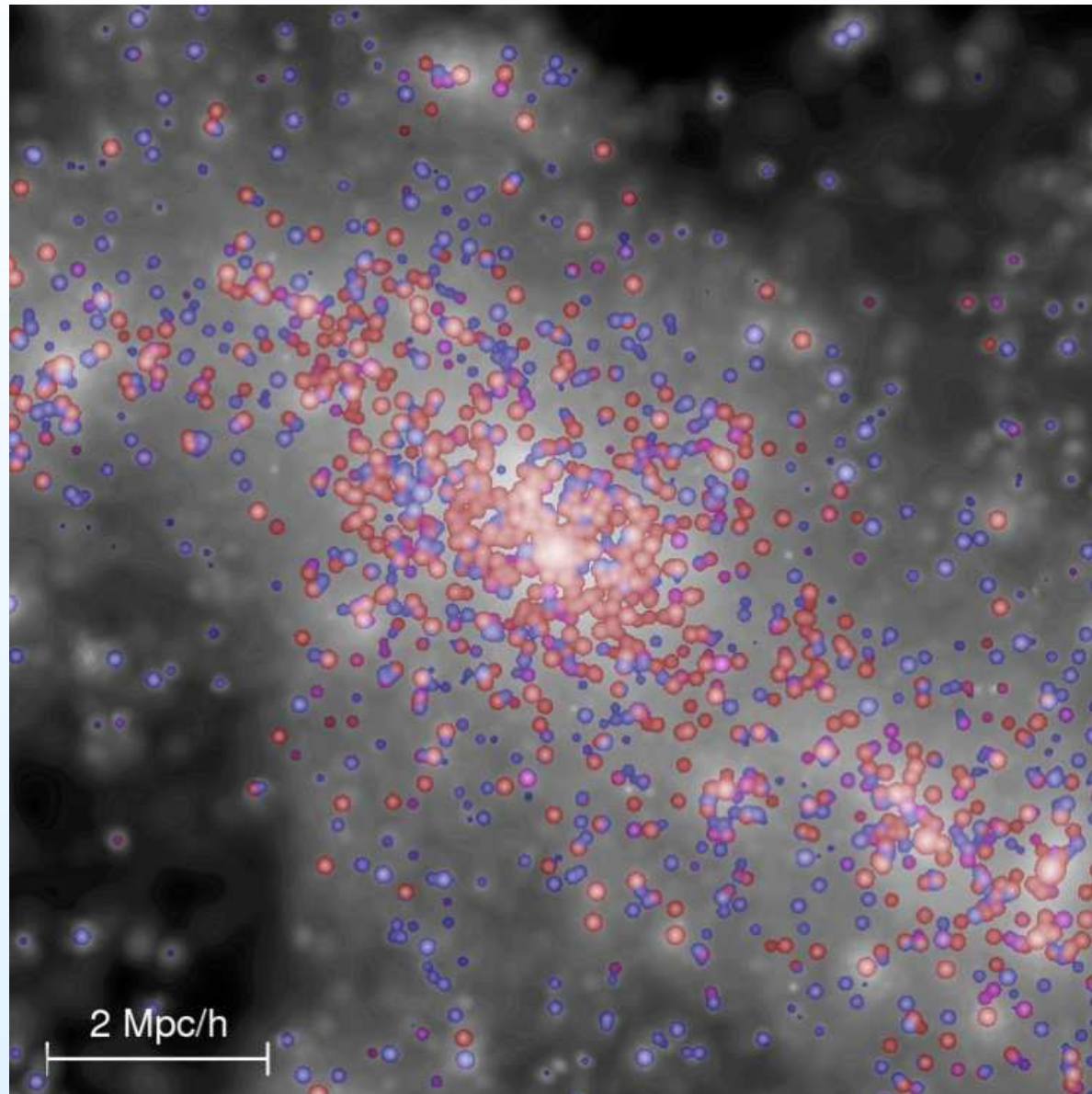
The Galform model produces lots of outputs, some of which can be compared to observed data from the real universe.

Questions of interest. Are there any choices of input parameters that will make Galform output resemble the observed universe? If so, how many different parameter choices are there?

The Dark Matter Simulation



The Galform Model



Inputs

To perform one run, we need to specify the following 17 inputs:

vhotdisk:	100 - 550	VCUT:	20 - 50
aReheat:	0.2 - 1.2	ZCUT:	6 - 9
alphacool:	0.2 - 1.2	alphastar:	-3.2 - -0.3
vhotburst:	100 - 550	tau0mrg:	0.8 - 2.7
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We choose 11 outputs that are representative of the Luminosity functions and emulate the functions $f_i(x)$.

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	No. Model Runs	No. Active Vars	Space Remaining
Wave 1	1000	5	14.9 %
Wave 2	1414	8	5.9 %
Wave 3	1620	8	1.6 %
Wave 4	2011	10	0.12 %

Linking models to reality: exchangeability

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The exchangeability representation for random functions allows us to express each function as

$$f^r(x) = M(x) \oplus R^r(x)$$

$M(x)$ is the mean function

$R^r(x)$ are the (uncorrelated, exchangeable, mean zero) residual functions.

Linking models to reality: exchangeability 2

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If each function is emulated as

$$f^r(x) = B^r g(x) + e^r(x),$$

then we have

$$B^r = M_B + R_B^r, e^r(x) = M_e(x) + R_e^r(x)$$

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If we consider dark matter in our universe to be exchangeable with the 520 individual simulations, then our emulator for Galform evaluated on the correct dark matter configuration is

$$f^*(x) = (M_B + R_B^*)g(x) + M_e(x) + R_e^*(x)$$

We cannot evaluate this simulator (because we don't know the appropriate dark matter configuration) but we can emulate it, based on a detailed analysis of the Galform experiments.

Linking models to reality: exchangeability and reification

This is an example of the idea of reification (from reify: to treat an abstract concept as if it were real), as follows.

The reason that the evaluations of the simulator are informative for the physical system is that the evaluations are informative about the general relationships between system properties, x , and system behaviour y .

Linking models to reality: exchangeability and reification

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The reason that the evaluations of the simulator are informative for the physical system is that the evaluations are informative about the general relationships between system properties, x , and system behaviour y .

More generally, evaluations of a collection of models are jointly informative for the physical system as they are jointly informative for these general relationships.

Linking models to reality: reification

Therefore, our inference from model to reality proceeds in two parts.

[1] Emulate the relationship between system properties and system behaviour (we call this the “reified model”).

[2] Decompose the difference between model and physical system as

[A] The difference between our simulator and the reified form.

[B] The difference between the reified form at the physically appropriate choice of x and the actual system behaviour y .

Linking models to reality: reification

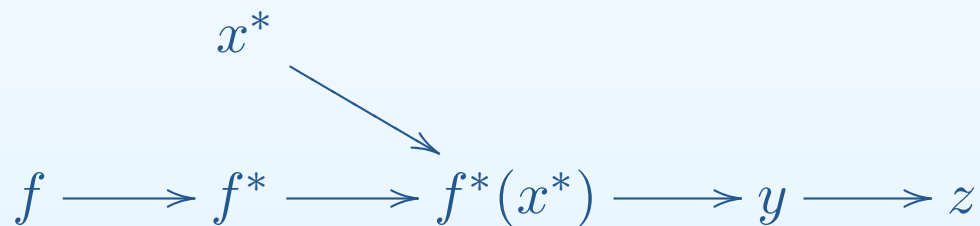
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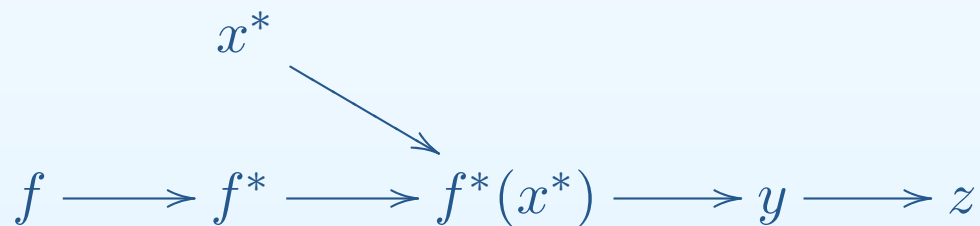
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Comment Reification retains tractability in linking computer evaluations and system values with system behaviour, while (i) removing logical problems in treatment of discrepancy, (ii) providing guidance for discrepancy modelling, (iii) showing how to combine multi-model information (through f^*).

A Reified influence diagram

$$\left[F_{h:[n]}^1(x), \dots, F_{h:[n]}^m(x) \right]$$

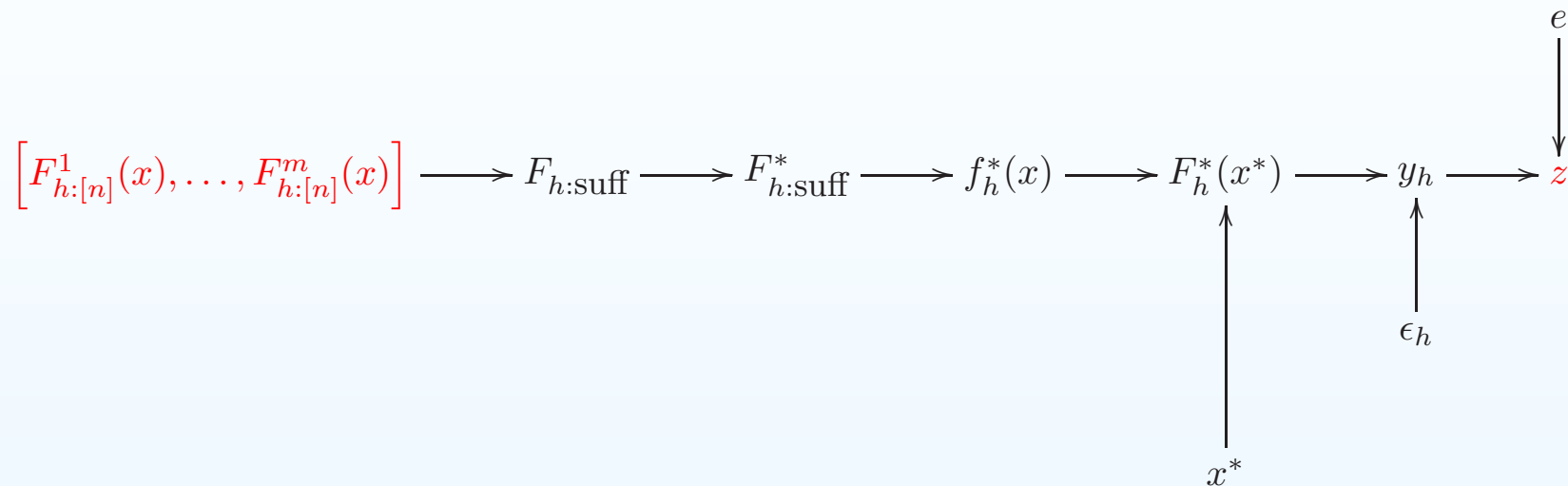
Evaluations of the simulator at each of m initial conditions
for historical components of simulator

A Reified influence diagram

$$\left[F_{h:[n]}^1(x), \dots, F_{h:[n]}^m(x) \right] \longrightarrow F_{h:\text{suff}} \longrightarrow F_{h:\text{suff}}^* \longrightarrow f_h^*(x)$$

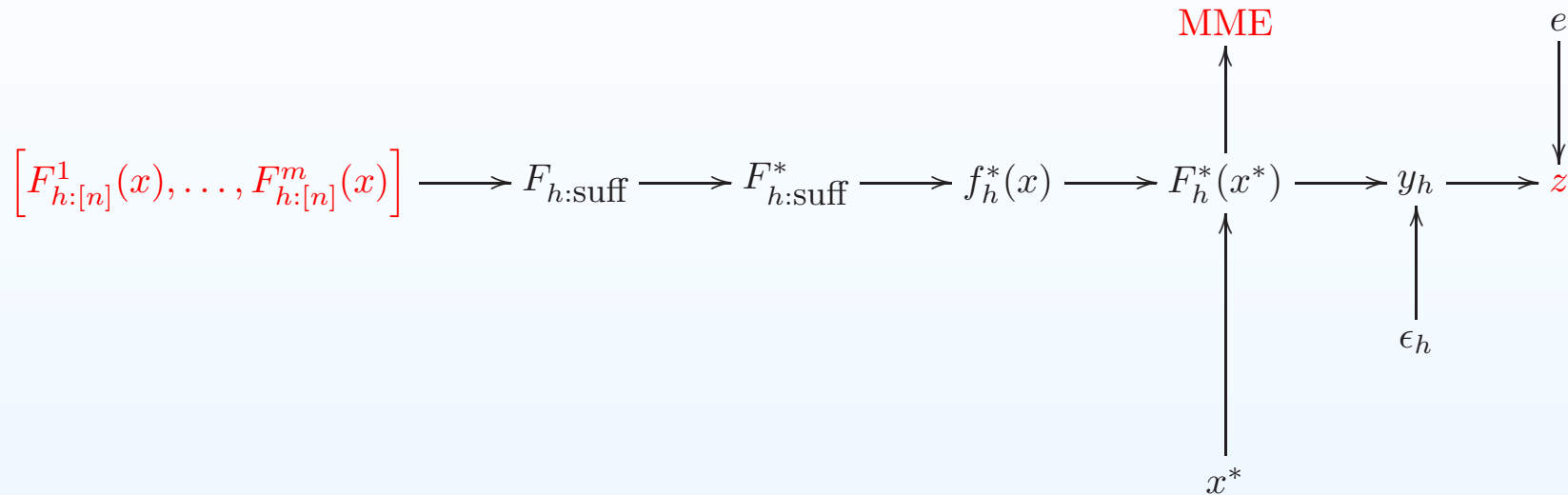
Global information $F_{h:\text{suff}}$ (from second order exchangeability modelling).
passes to Reified global form and to reified emulator.

A Reified influence diagram



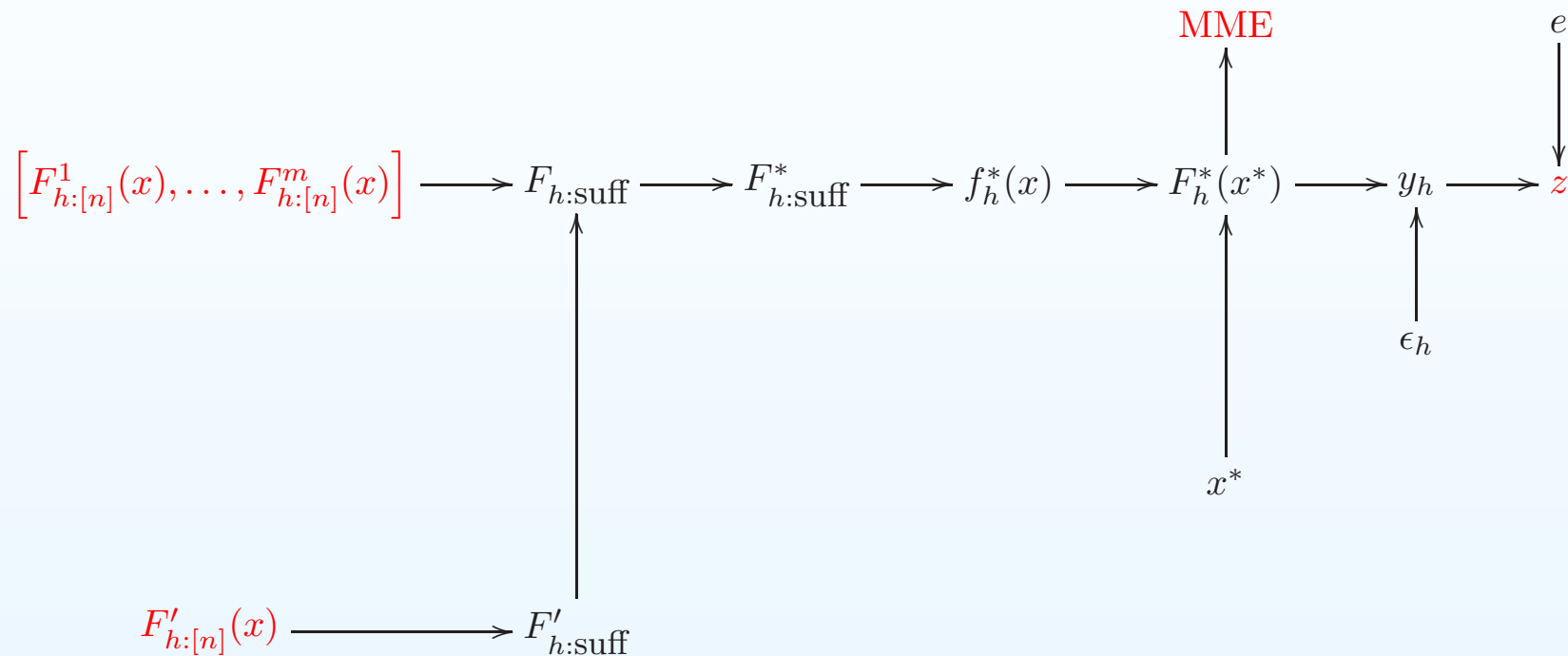
Link with x^* to reified function, at true initial condition, linked to data z

A Reified influence diagram



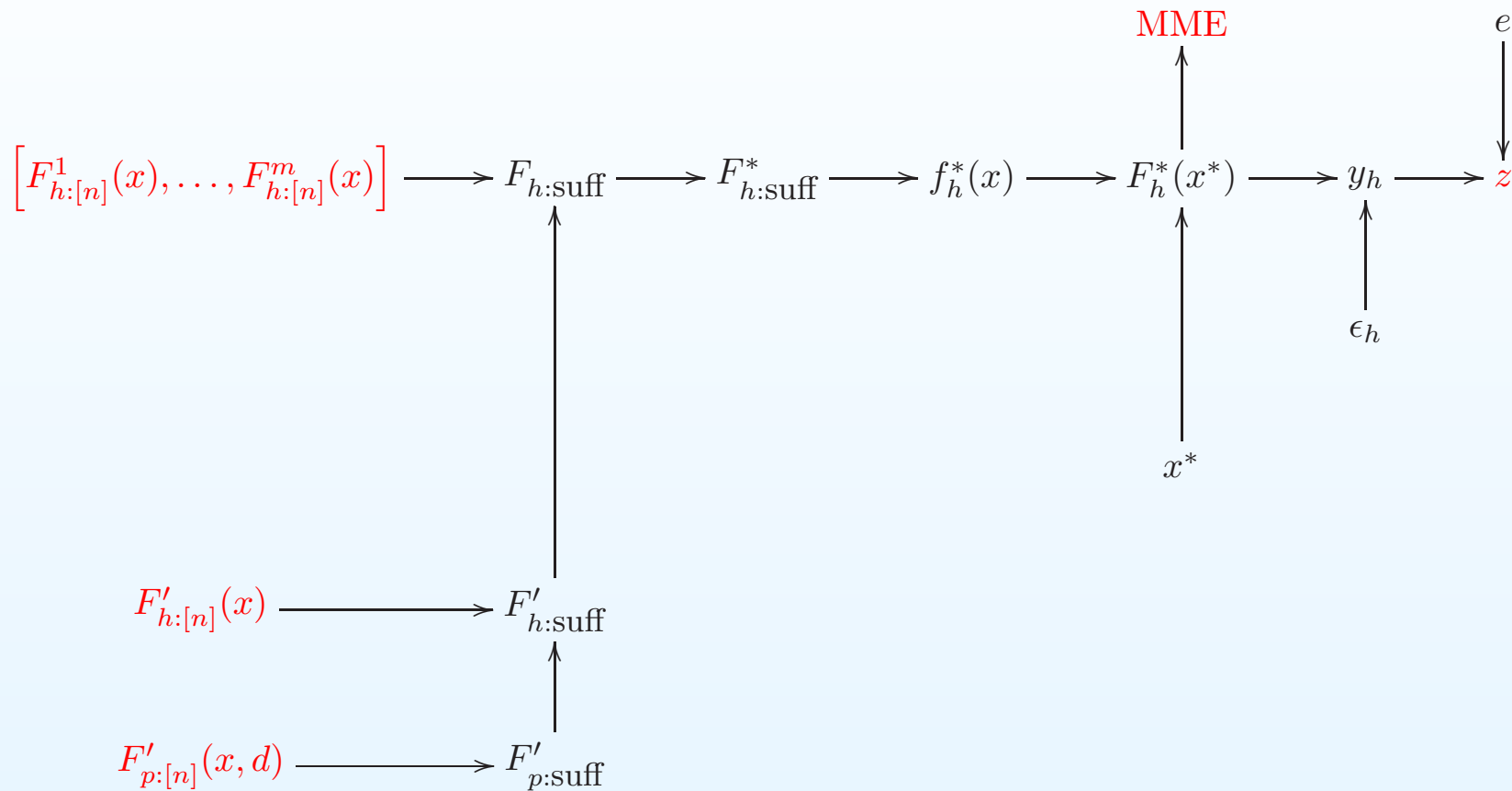
Add observation of a related multi-model ensemble (MME) consisting of tuned runs from related models (more exchangeability modelling).

A Reified influence diagram



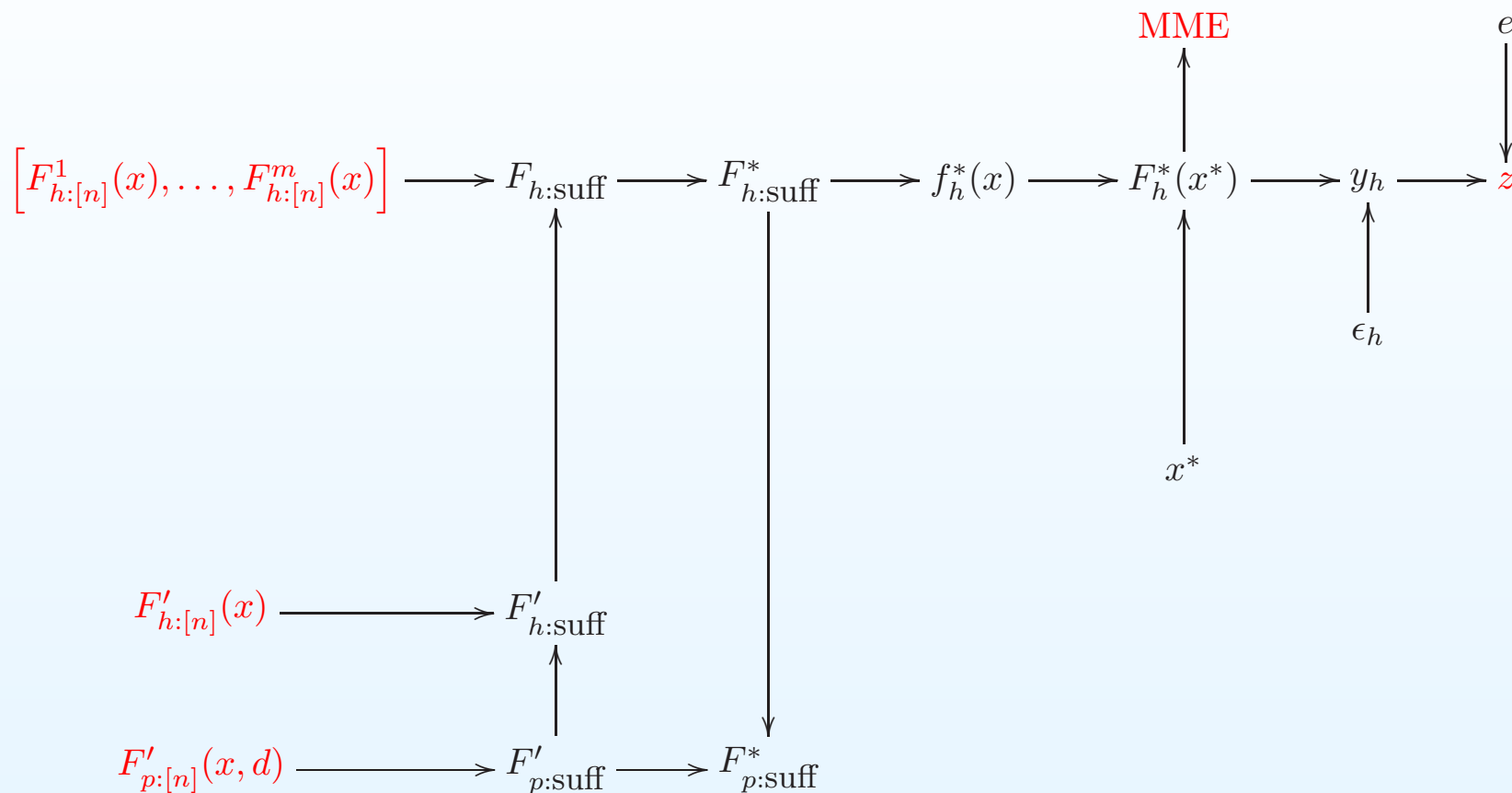
Add a set of evaluations from a fast approximation

A Reified influence diagram



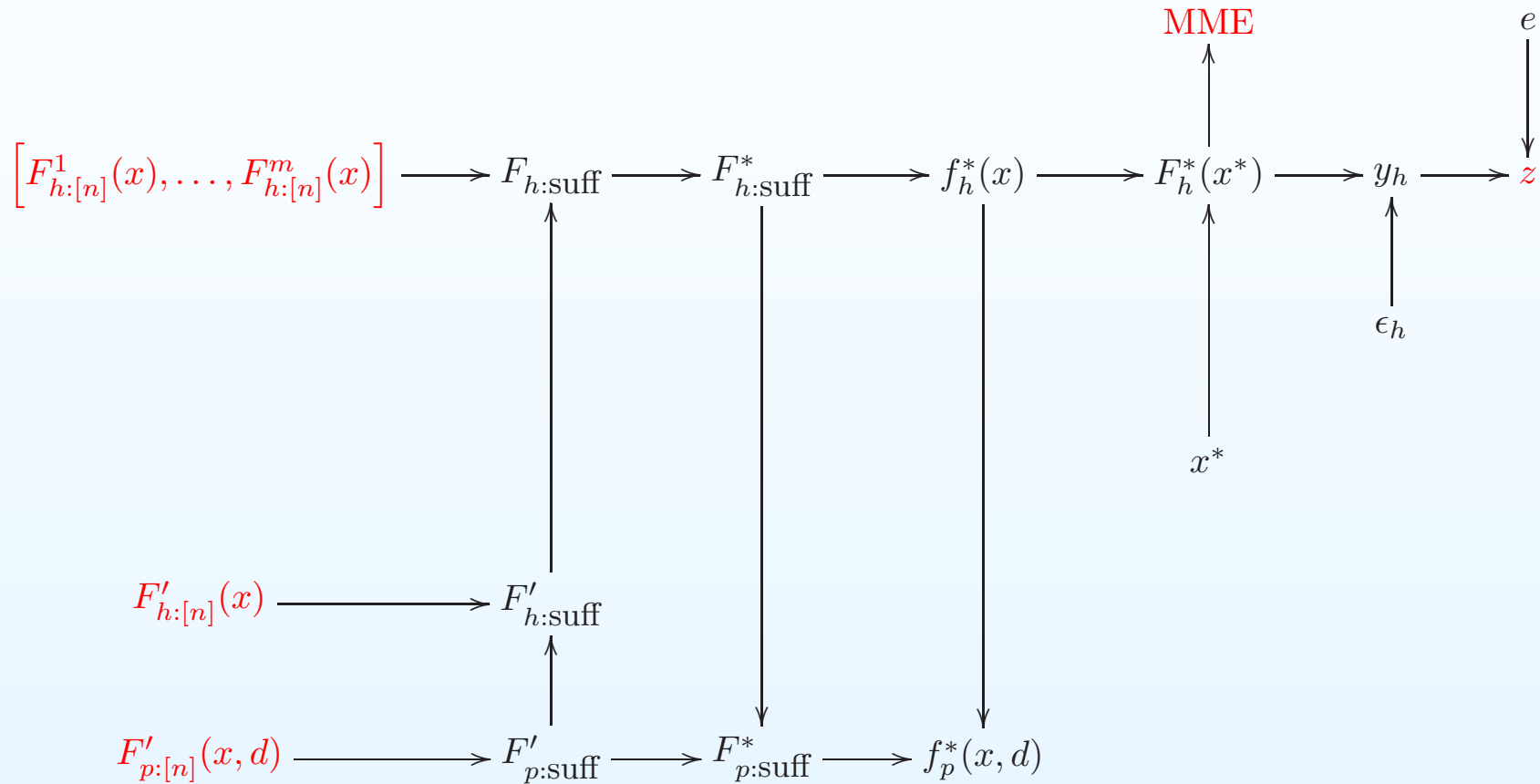
Add evaluations of fast simulator for outcomes to be predicted, with decision choices d

A Reified influence diagram



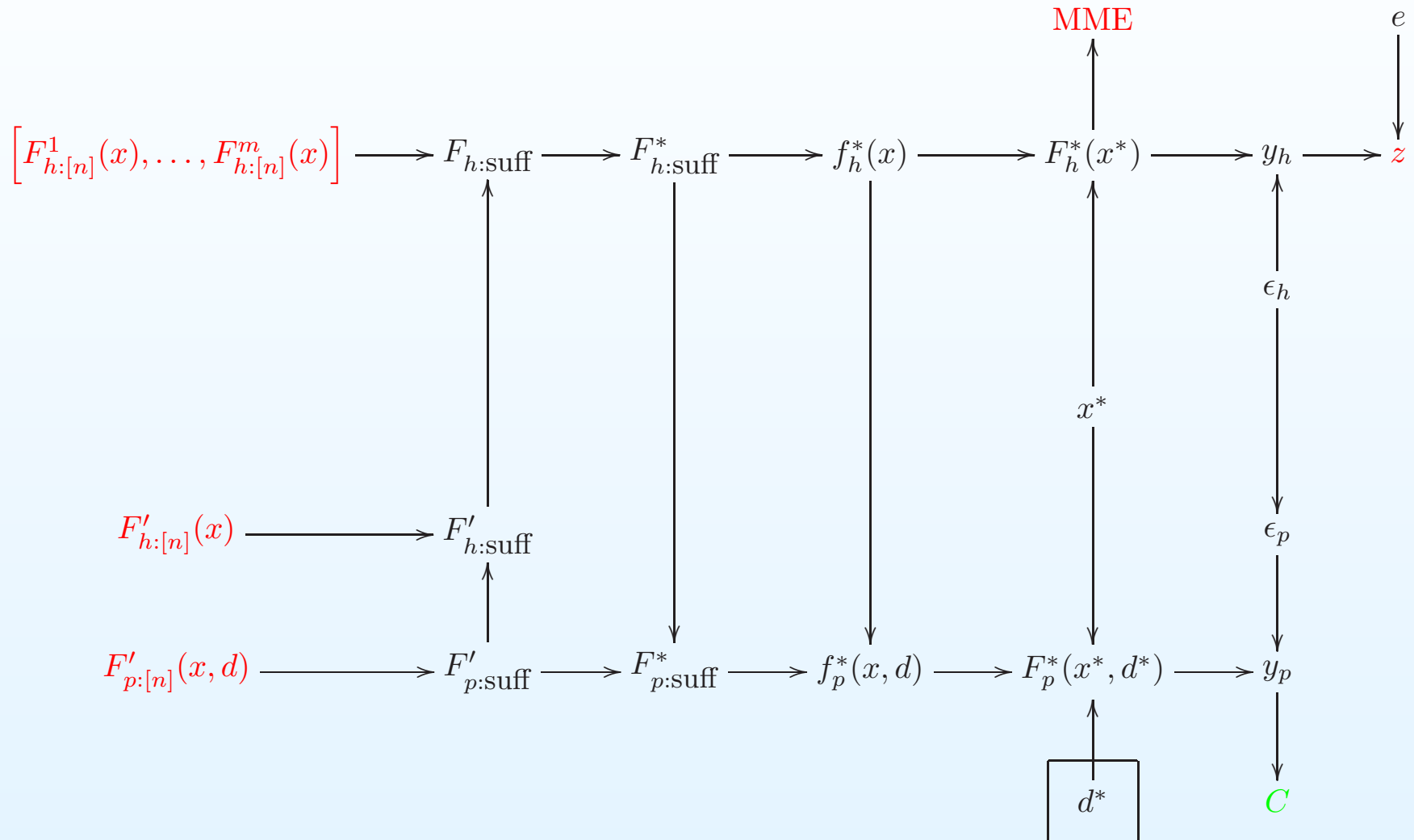
Link to reified global terms for quantities to be predicted

A Reified influence diagram



And to reified global emulator, based on inputs and decisions

A Reified influence diagram



And link, through true future values y_p , to the overall utility cost C of making decision choice d^* [Attach more models to diagram at $F^*(x^*)$]

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The main problems are cultural - the lack of people, resources, and will to recognise the urgency and importance of the task
(but we could change that).

References for applications

Reservoir Application

P.S. Craig, M. Goldstein, A.H. Seheult, J.A. Smith (1997). Pressure matching for hydrocarbon reservoirs: a case study in the use of Bayes linear strategies for large computer experiments (with discussion), in *Case Studies in Bayesian Statistics*, vol. III, eds. C. Gastonis et al. 37-93. Springer-Verlag.

J. Cumming, M. Goldstein Bayes Linear Uncertainty Analysis for Oil Reservoirs Based on Multiscale Computer Experiments (2009), in the *Handbook of Applied Bayesian Analysis*, eds A. O'Hagan, M. West, OUP

Galaxy Application

I. Vernon, M. Goldstein, R. Bower (2010), “*Galaxy Formation: a Bayesian Uncertainty Analysis* (with discussion) ”, *Bayesian Analysis*, 5(4): 619–670.

Climate Application

D. Williamson, M. Goldstein, A. Blaker (2012) Fast Linked Analyses for Scenario based Hierarchies, *JRSSC*, 61, 665-691.

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M. Goldstein Subjective Bayesian analysis: principles and practice (2006)
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